LOW NOISE SIS MIXER FOR THE FREQUENCY ABOVE 1 THz

Alexander Karpov, D. Miller, J. Zmuidzinas California Institute of Technology, 1200 E. California Blvd., Pasadena, CA 91125

J.A. Stern, B. Bumble, H.G. LeDuc Jet Propulsion Laboratory, 4800 Oak Grove Drive, Pasadena, CA 91109

ABSTRACT

We developed a SIS mixer for the 1.1 - 1.25 THz band of the heterodyne receiver of the Herschel space observatory. Our approach may be used up to 1.6 THz. The quasi-optical SIS mixer has two NbN/AlN/Nb junctions with the critical current $30\text{-}50~\text{kA/cm}^2$ and the gap voltage of 3.4~mV. The tuning circuit integrated with SIS junction has the base electrode of Nb and a gold wire layer. This approach simplifies the SIS junction technology, compared to a design using NbTiN base electrode. The junction base electrode and the ground of the tuning micro strip circuit are formed in one step. The frequency of operation of the mixer is well above the gap frequency of Nb, and it behaves here as a normal metal. The resistivity of Nb at the critical temperature of $0.2~\text{m}\Omega$ cm is below the resistivity of the best normal metal films. The measured receiver noise temperature is below 600~K.

Editor's note: The manuscript of this presentation is not available for the proceedings.

Contact information for A. Karpov: Email: karpov@submm.caltech.edu